Discovery space and science with the PLACID stellar coronagraph

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Introduction

 PLACID (Programmable Liquid-crystal Active Coronagraph Imager for the DAG telescope) is a coronagraphic instrument, providing adaptive high-contrast imaging capbilities from H - to Ks - band

• PLACID uses a pixelated Spatial Light Modulator (SLM) to generate coronagraphic focal-plane masks (FPMs) for the first time on a telescope (4 m, DAG observatory, Erzurum, Turkey)

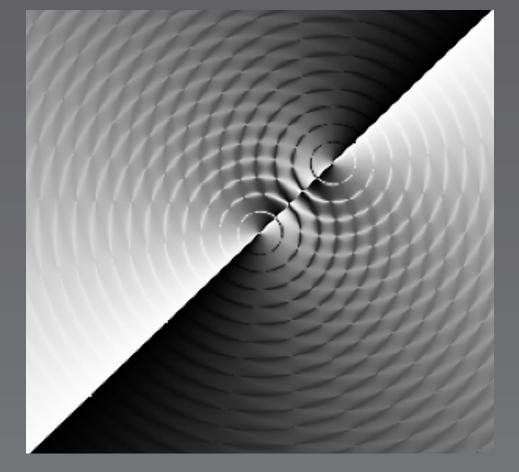
• Instrument delivered in March '24, first light expected by end of 2024

• Remote reconfiguration on-demand to adapt to e.g. observing





conditions, multiple star coronagraphy, correcting aberrations, segmented primary mirrors (ELTs, HWO, ...)



Vortex FPM (n=2) programmed for a binary star (can be rotated in time to perform ADI) [1]



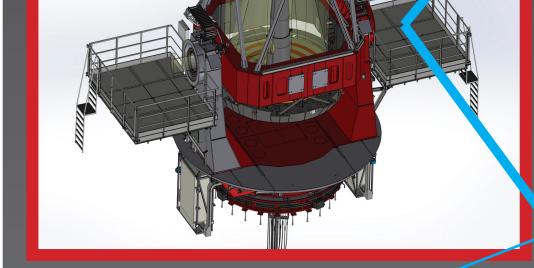
 10^{-1}



Known Directly Imaged Exoplanets/Candidates

Exoplanet, $t_{exp} \geq 8h$

TROIA O expected performance TROIA: Strehl ratio vs. Vmag





PLACID: active (SLM-based) coronagraph instrument jointly developed at University of Bern and HEIG-VD Yverdon, Switzerland

PLACID targets and discovery space

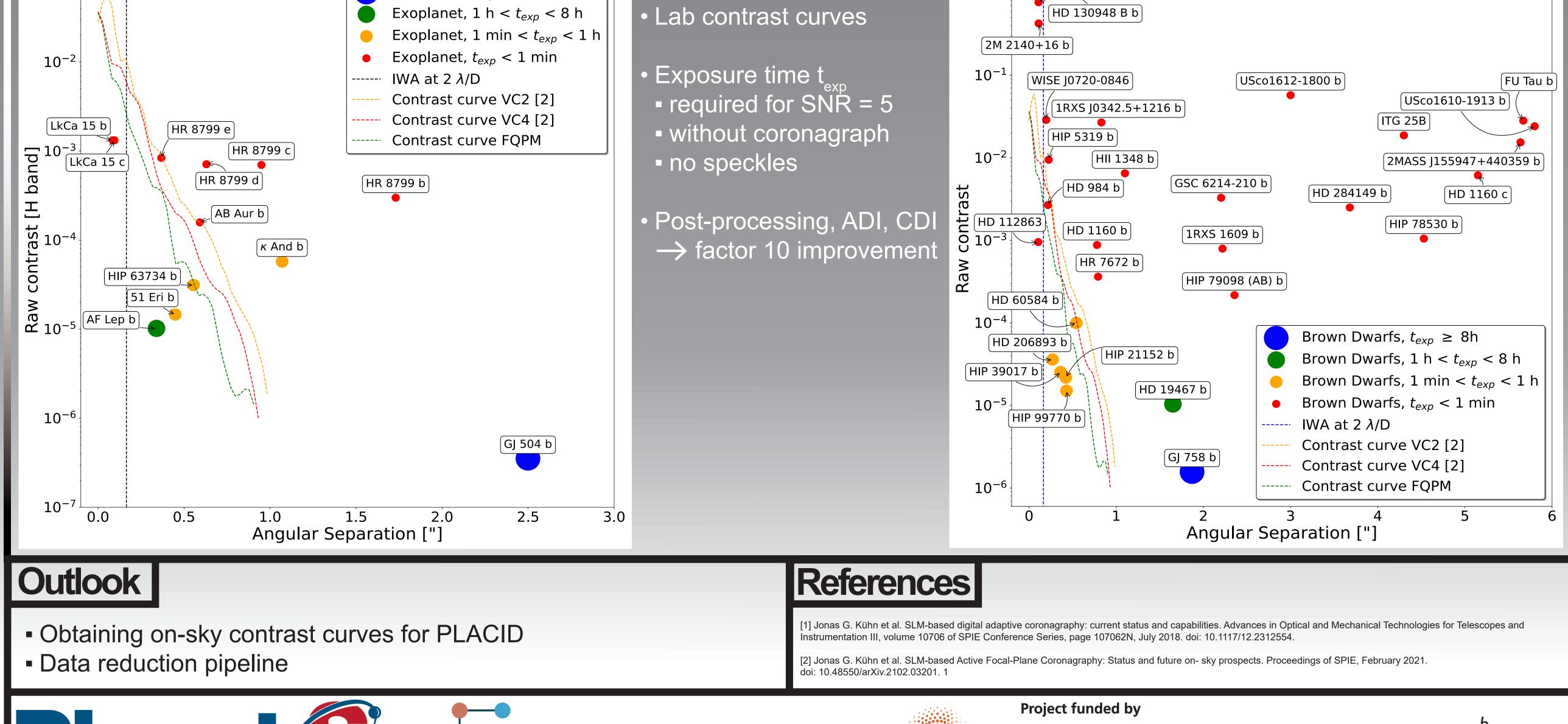
 Known exoplanets / candidates, Brown Dwarfs, circumstellar disks, binaries/triples

• Gaia, TESS, PLATO direct-imaging follow-ups in the North

PLACID observational constraints:
Site: DEC: ≥ -24°
TROIA AO guide star: V ≤ 13 mag
On-sky FOV: 16" x 9.6"

Known Directly Imaged Brown Dwarfs

10⁰



• Plots: no sky degradation

